

Simulations of Cosmic Ray Interactions: Past, Present, and Future

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Physics at the End of the Galactic Cosmic Ray Spectrum

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Outline:

- 1) Historical trend
- 2) High energy interactions - basic approach
- 3) Matching with QCD?
- 4) Non-linear effects
- 5) Air shower predictions
- 5) Model uncertainties
- 6) Future prospects

Historical Trend

Simulations of CR interactions - 50 year history

(G.Zatsepin, M.Hillas, M.Dedenko, and others)

Early approaches - semi-inclusive:

- sampling of K^{inel} , $N_{\pi^{\pm}}$, N_{π^0}
- simple parameterizations of inclusive spectra for given K^{inel} , $N_{\pi^{\pm}}$, N_{π^0}

Microscopic MC models (around 90s):

requested by the increased accuracy of CR experiments

Evolution of physics picture:

- first simple scaling: $\sigma_{\text{inel}}(s) = \text{const}$, $K^{\text{inel}}(s) = \text{const}$, $N_{\pi}(s) = \text{const}$
- but $\sigma_{\text{inel}}(s)$ - increasing with s
- $N_{\pi}(s)$ - also increasing
- finally, the inelasticity grows with s

Presently: [phenomenological MC models](#)

[High energy models](#) ($> 10 \div 100$ GeV):

- DPMJET (Engel, Ranft, Roesler)
- neXus (Drescher, Hladik, SO, Pierog, Werner)
- QGSJET (Kalmykov, SO)
- SIBYLL (Engel, Gaisser, Lipari, Stanev)

[Low-energy models](#):

- GHEISHA/GEANT (Fesefeldt)
- Splitting algorithm (Hillas)
- FLUKA (Fasso, Ferrari, Ranft, Sala)
- UrQMD (Bass, Bleicher et al.)
- TARGET (Engel, Gaisser, Protheroe, Stanev)
- HADRIN/NUCRIN (Hanssgen, Ranft)

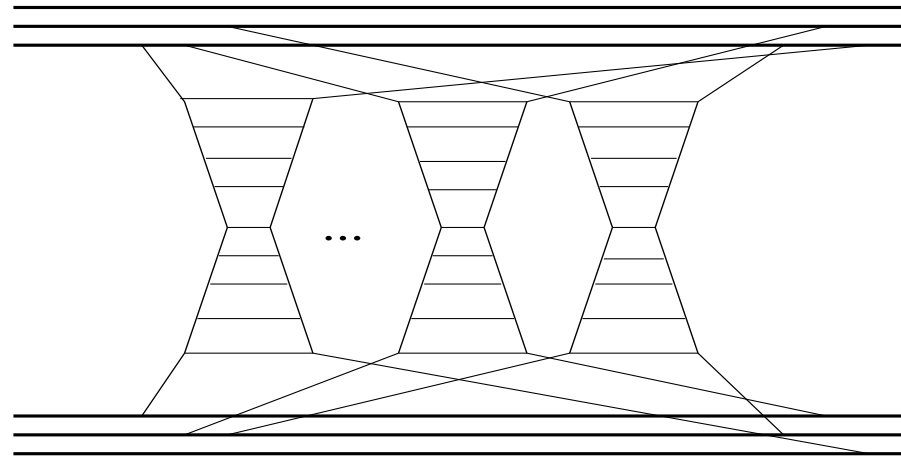
Current and future progress - [due to closer link to QCD](#)

But: (most likely) the models will [remain phenomenological](#)

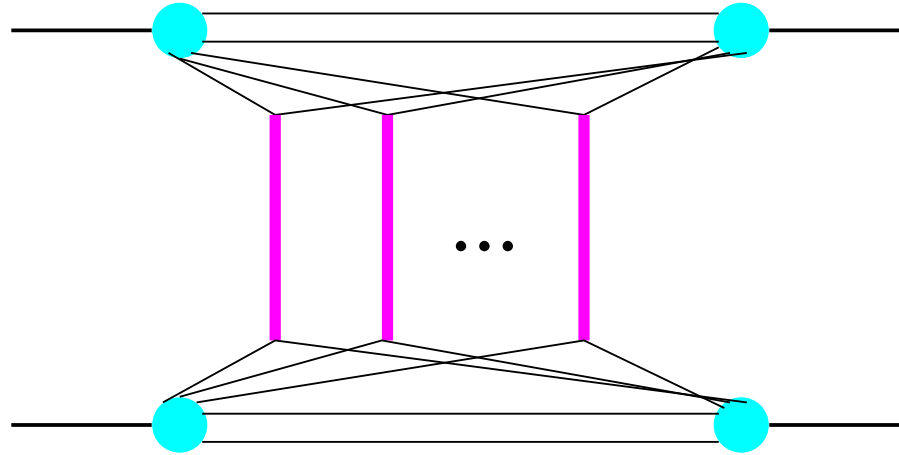
Basic Approach

High energy interactions - multiple scattering scheme:

- required by unitarization
- inclusive cross sections increase faster than total ones
- describes diffraction dissociation



Elementary interaction = parton cascade \equiv Pomeron exchange



Pomeron amplitude:

$$f_{ab}^{\text{P}}(s, b) = \frac{i\gamma_a\gamma_b s^{\alpha_{\text{P}}(0)}}{\lambda_{ab}(s)} \exp\left(\frac{-b^2}{4\lambda_{ab}(s)}\right) \equiv i\chi_{ab}^{\text{P}}(s, b)$$

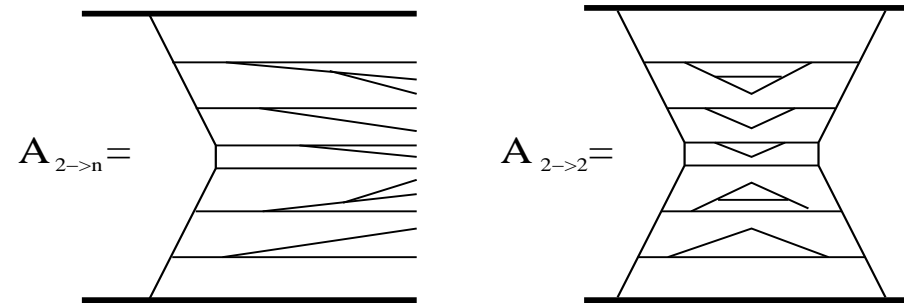
$$\lambda_{ab}(s) = R_a^2 + R_b^2 + \alpha'_{\text{P}}(0) \ln s$$

Pomeron intercept $\alpha_{\text{P}}(0) > 1$ - energy increase

Pomeron slope $\alpha'_{\text{P}}(0)$ - increasing spacial size of the interaction

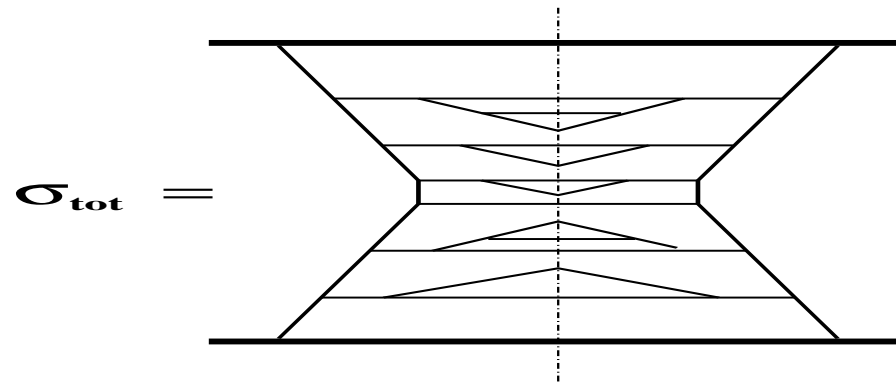
Cross sections & particle production

Elementary interaction - inelastic & elastic amplitudes:



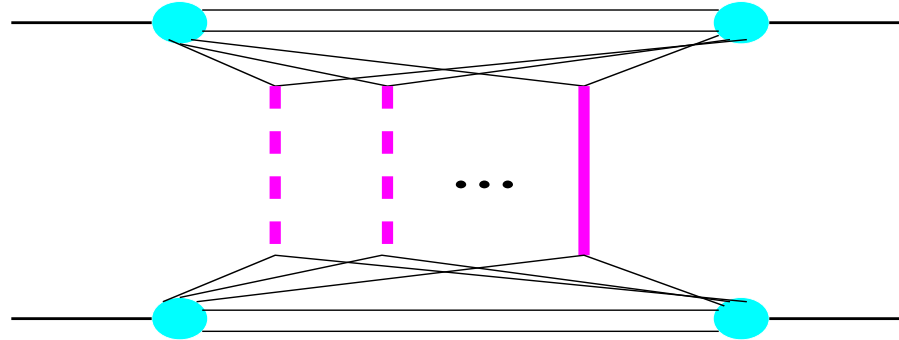
Cross section - **optical theorem**:

$$\sigma_{\text{tot}} = \sum_n \int d\tau_n A_{2 \to n} \cdot A_{2 \to n}^* = 2\text{Im} A_{2 \to 2}$$



AGK cutting rules (V.Abramovskii, V.Gribov, O.Kancheli, 1973):

no interference between different classes of the interaction \Rightarrow cross sections:



$$\sigma_{ab}^{\text{tot}}(s) = 2 \sum_{i,j} C_i^a C_j^b \int d^2b \left[1 - e^{-\lambda_i^a \lambda_j^b \chi_{ab}^{\text{P}}(s,b)} \right]$$

$$\sigma_{ab}^{\text{inel}}(s) = \sum_{i,j} C_i^a C_j^b \int d^2b \left[1 - e^{-2\lambda_i^a \lambda_j^b \chi_{ab}^{\text{P}}(s,b)} \right]$$

“Topological” cross section (n “cut” Pomerons):

$$\sigma_{ab}^{(n)}(s) = \sum_{i,j} C_i^a C_j^b \int d^2b \frac{\left[2\lambda_i^a \lambda_j^b \chi_{ab}^{\text{P}}(s,b) \right]^n}{n!} e^{-2\lambda_i^a \lambda_j^b \chi_{ab}^{\text{P}}(s,b)}$$

C_i^a, λ_i^a - relative weights and strengths of diffraction eigenstates of hadron a

Important: diffraction and inelastic screening - related to each other:
higher diffraction \Rightarrow larger inelastic screening
 \Rightarrow smaller cross sections at higher energies and on nuclei

Dual Parton Model (A.Capella et al., 1979),

Quark-Gluon String model (A.Kaidalov, K.Ter-Martyrosyan, 1982):

“cut” Pomeron \Rightarrow string formation & break up \Rightarrow hadronization

Remarkably simple picture:

- elementary processes - Pomeron exchanges
- small number of adjustable parameters
- generalization to hA , AA - parameter free

Experimental data on σ^{tot} , σ^{el} , σ^{difr} , B^{el} \Rightarrow fixing Pomeron parameters

Experimental data on particle production \Rightarrow fixing hadronization parameters

Drawbacks:

- high energy behavior governed by the Pomeron intercept, not by QCD
- no space for high p_t physics

Matching with QCD?

Present approaches:

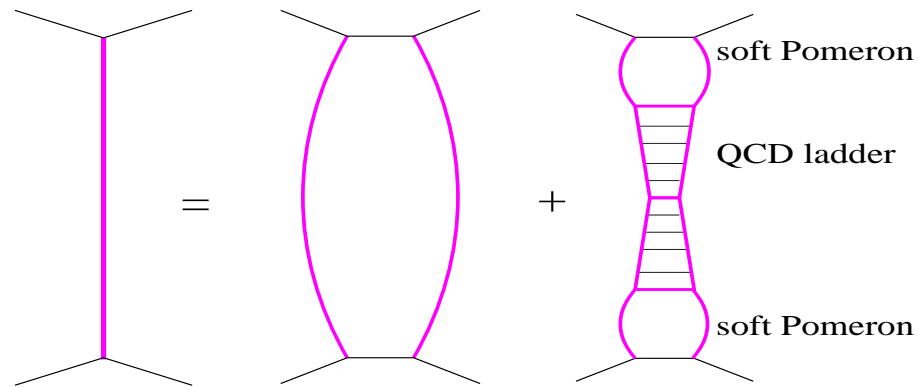
- “mini-jet” scheme - SIBYLL model (R.Fletcher et al., 1994; R.Engel et al., 1999), PYTHIA model (T.Sjostrand, 1987), others
- “semi-hard Pomeron” scheme - QGSJET model (N.Kalmykov, SO, A.Pavlov, 1994, 1997), neXus model (H.Drescher et al., 2001)

“Semi-hard Pomeron” approach:

- introduce a “threshold” scale Q_0^2
- use “soft” Pomeron below Q_0^2
- use DGLAP ladder for $|q^2| > Q_0^2$

The same picture as before but based on a “general Pomeron”:

$$\chi_{ab}^P(s, b) = \chi_{ab}^{P_{\text{soft}}}(s, b) + \chi_{ab}^{P_{\text{sh}}}(s, b)$$



“Mini-jet” approach - without “soft” Pomeron (no “soft pre-evolution”)

⇒ implies screening mechanism: $Q_0^2 = Q_0^2(s)$

(to suppress non-perturbative parton evolution)

In general, Pomeron physics is complicated:

- QCD parton evolution
- string formation and break-up
- fragmentation into hadrons

Interaction physics - **simple and transparent**:

- a number of elementary processes - Pomeron (QCD ladder) exchanges
- small number of adjustable parameters
- generalization to hA , AA - parameter free
- high energy asymptotics - governed by pQCD

Important: **elementary processes proceed independently**

Rapidly comes to its limits - cross sections and hadron multiplicities increase too fast **for realistic parton structure functions** (SFs)

Non-Linear Effects

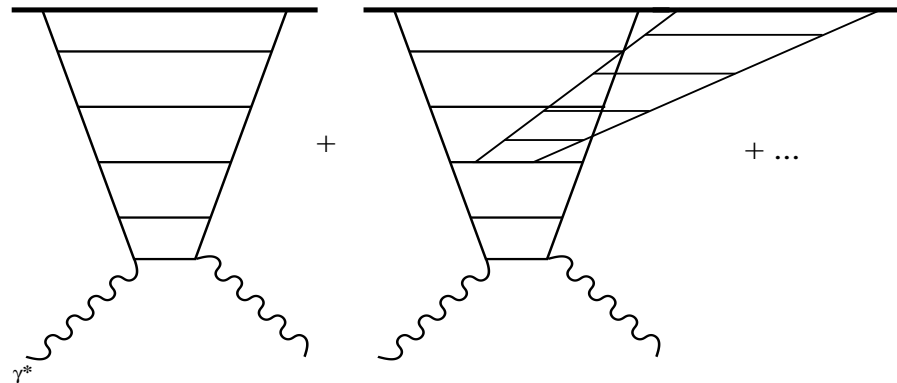
Independent interaction picture is **inadequate** for large s , small b , large A :

- many partons **closely packed**
- \Rightarrow **expected to interact with each other**

QCD approach (L.Gribov, E.Levin, M.Ryskin, 1983) - asymptotic picture:

- **parton saturation** at some scale $Q_{\text{sat}}^2(x) \Rightarrow$ “soft” contribution suppressed
- **QCD parton dynamics** for $Q_i^2 > Q_{\text{sat}}^2(x)$
- non-linear effects - **interaction between QCD ladders**

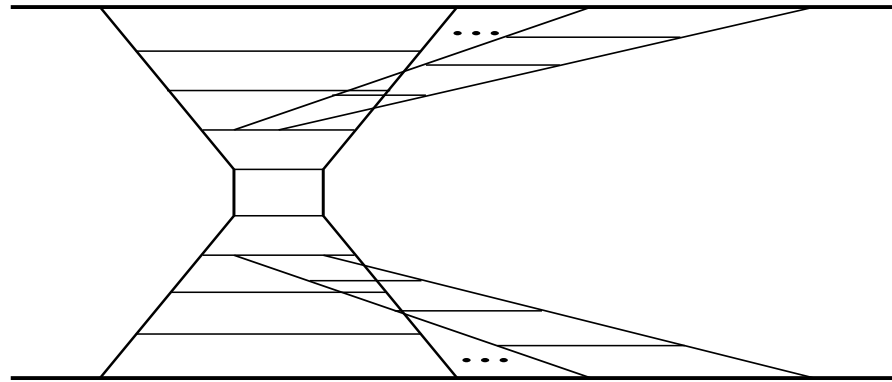
Structure function (SF) $F_2(x)$, $x \rightarrow 0$:



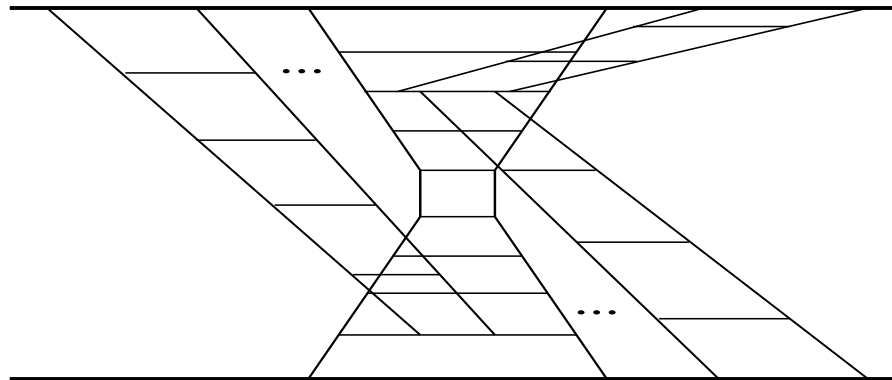
Provides formal justification for the “mini-jet” approach:

saturation-based picture; $Q_0^2 = Q_0^2(s)$ - effective saturation scale

Factorizable contribution to the eikonal - convolution of screened parton SFs:



But also non-factorizable contributions - impossible to account for:



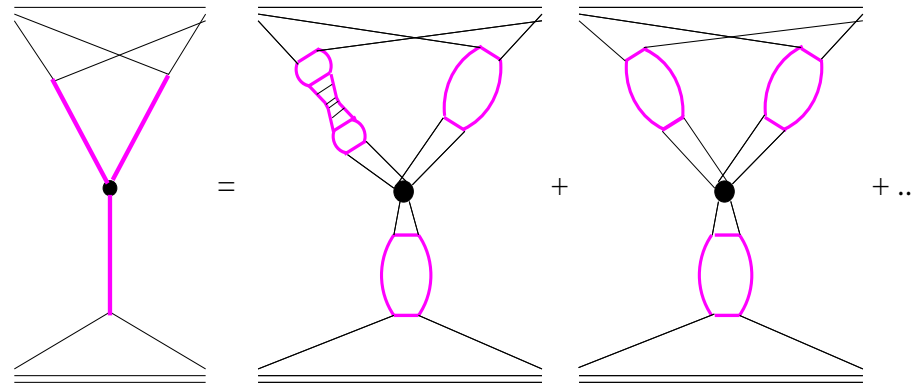
Does not account for:

- no saturation in peripheral interactions
- saturation being different in $hh-$, $hA-$, $AA-$ collisions
- screening effects in non-saturation regime

⇒ Treatment of realistic (not asymptotic) conditions needed!

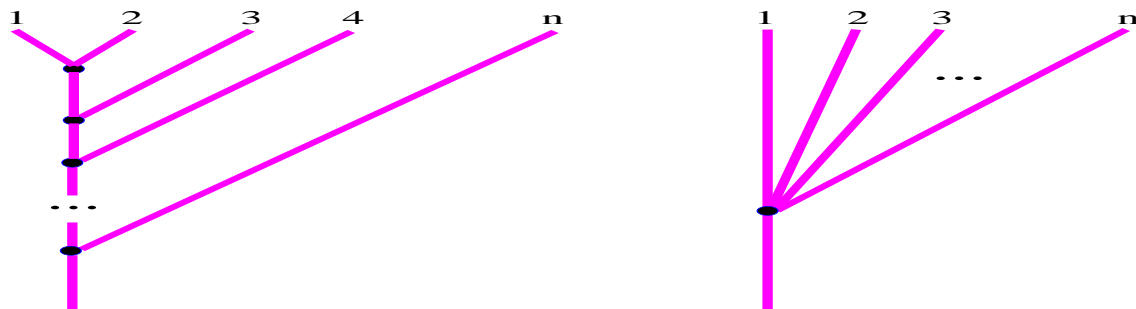
QGSJET-II model (SO, 2004) - alternative approach:

- assumes no saturation at a fixed Q_0^2 scale
- ⇒ non-linear effects = interactions between “soft” Pomerons
- ⇒ can be described using Gribov’s Reggeon scheme



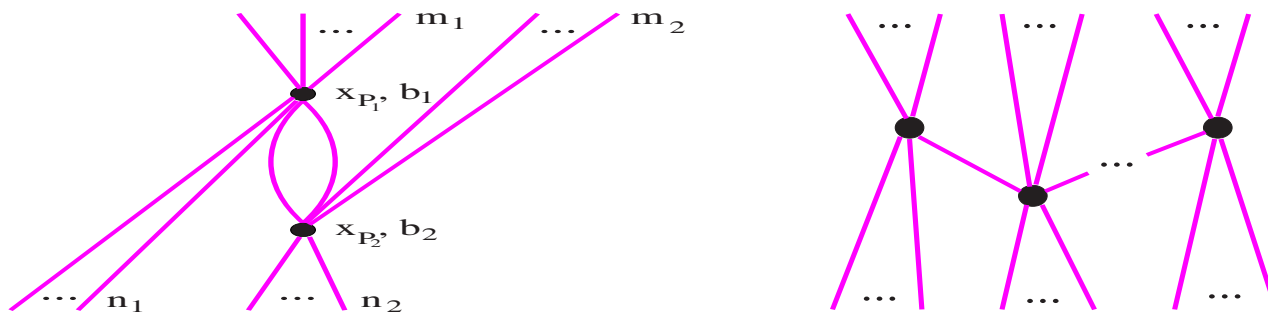
Main problem: all orders are to be accounted for!

Example: “fan” contributions with n “legs”:



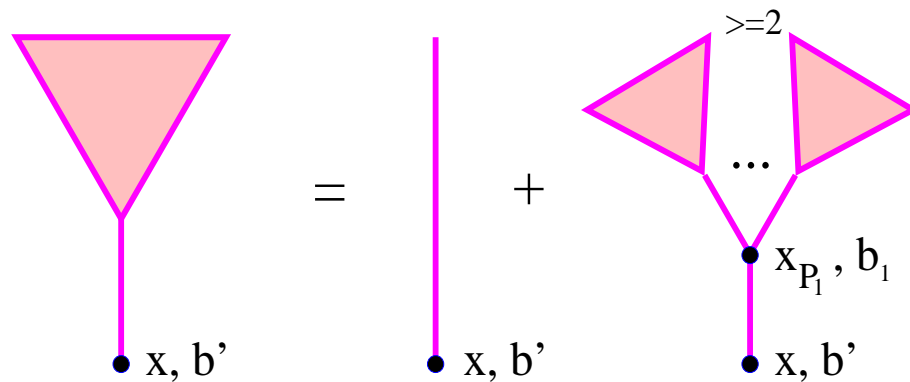
$$\Delta\chi_{ab}^{\text{fan}} = \sum_{n=1}^{\infty} \Delta\chi_{ab}^{\text{fan}} \sim \sum_{n=1}^{\infty} r_{3P}^n [-\chi_P(s)]^{n+1}, \quad \chi_P(s) \sim s^{\alpha_P(0)-1}$$

Still one can neglect “loop” and “chess-board” diagrams:

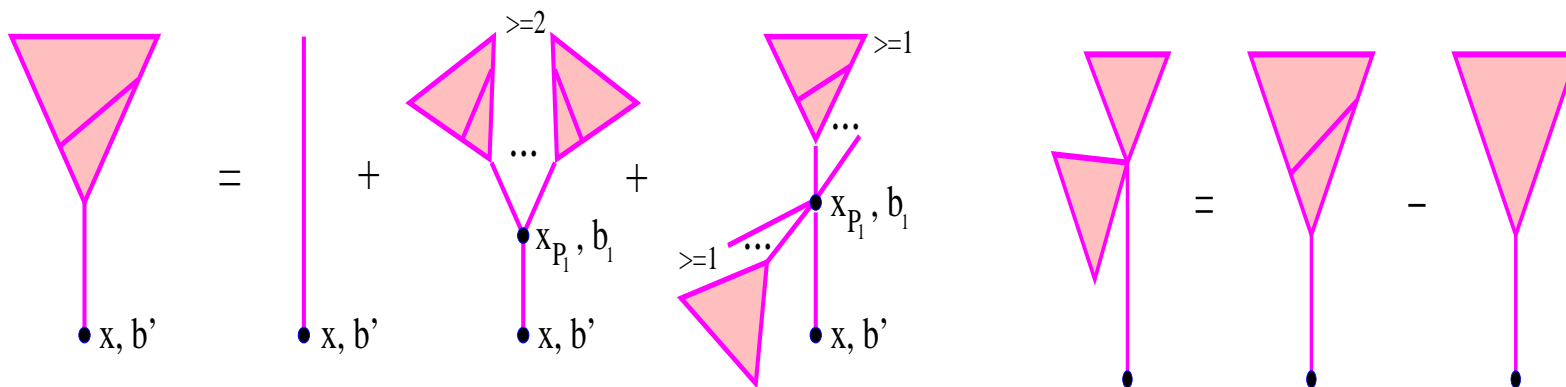


And re-sum all others ...

Elementary building blocks - “fan” contributions:



Introduce “generalized fans” and the difference:



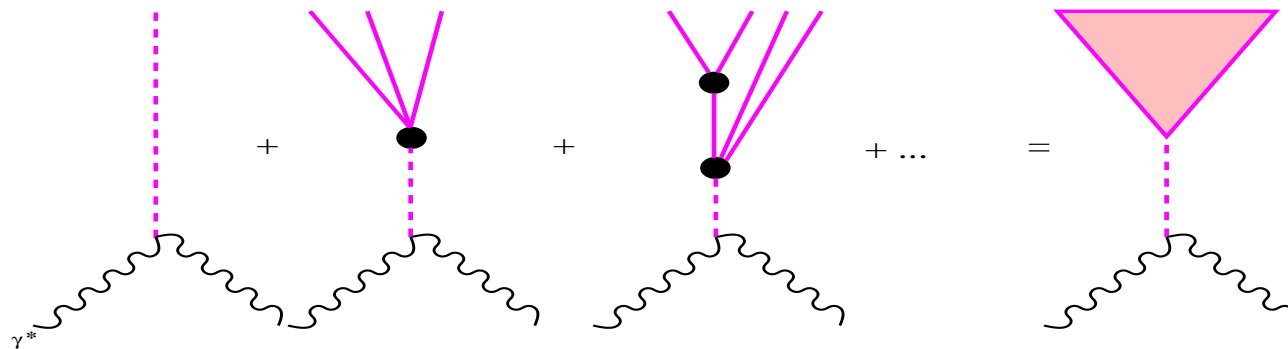
Now total eikonal:

$$\chi_{\text{tot}} = \left[\text{Diagram 1} + \text{Diagram 2} - \text{Diagram 3} \right] - \frac{1}{2} \left[\text{Diagram 4} - \text{Diagram 5} + \text{Diagram 6} - \dots \right]$$

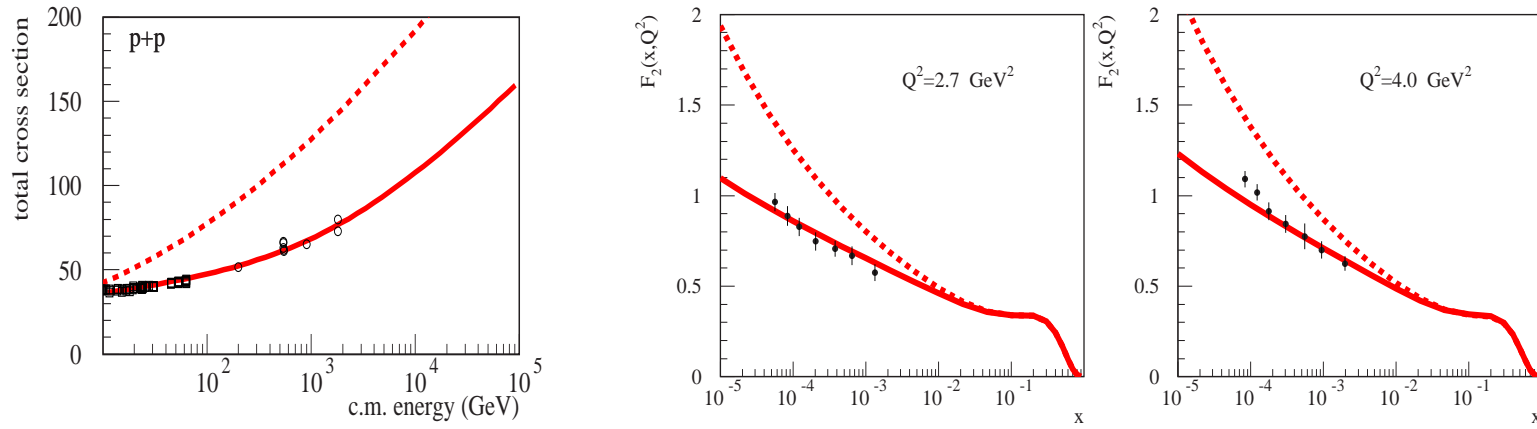
hA , AA - similar; “fans” include connections to different nucleons
 $\Rightarrow A$ -enhancement of screening effects

Cross sections - the same formulas as before but with the full eikonal $\chi_{ab}^{\text{tot}}(s, b)$

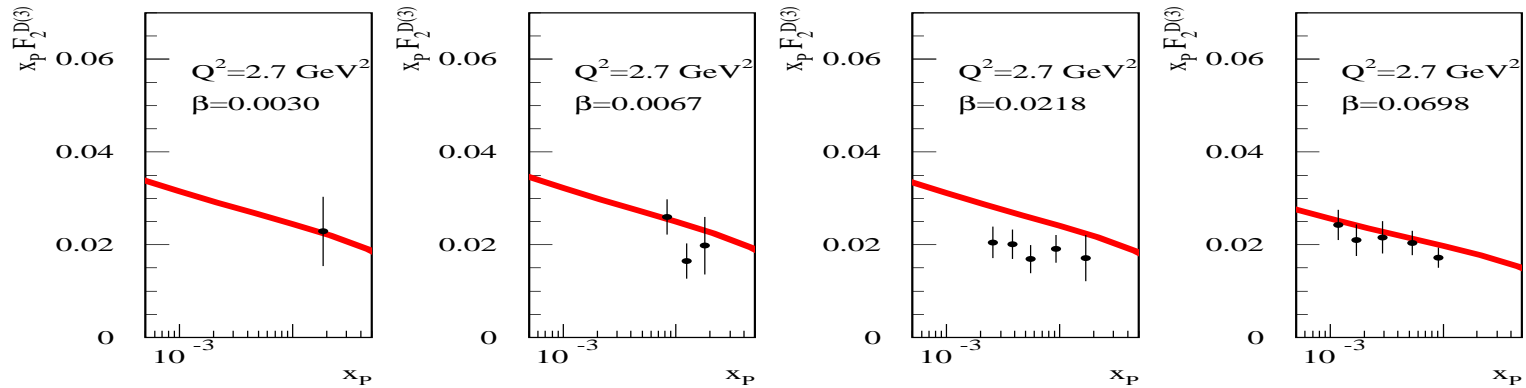
Proton SF $F_2(x, Q^2)$, $x \rightarrow 0$ - just a “fan”:



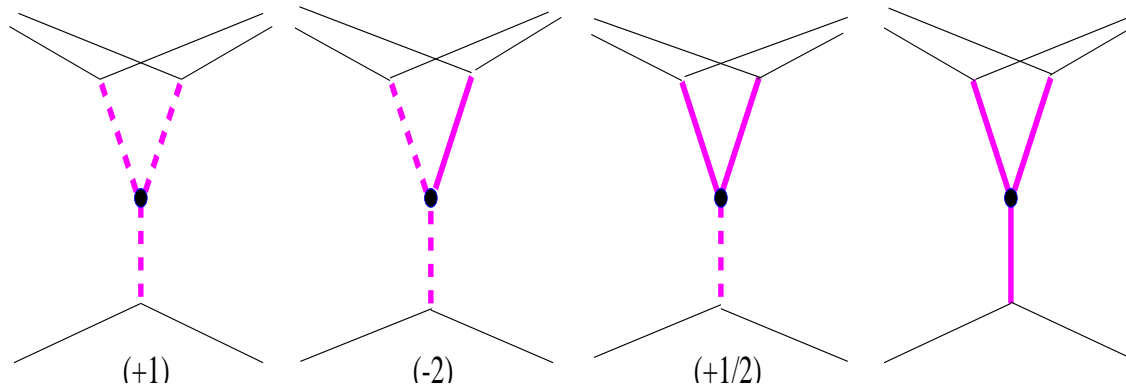
Total cross section and SF $F_2(x, Q^2)$ with (without) enhanced graphs
 ($Q_0^2 = 2.5 \text{ GeV}^2$, $\alpha_P(0) = 0.18$):



Triple-Pomeron vertex - can be inferred from “hard” diffraction data at HERA:



Not yet the end - one needs “cut” diagrams for particle production:



Total “cut” contribution can be obtained similarly

Finally we are back: physics of the interaction - (almost) simple and transparent

- a number of elementary processes (Pomeron + multi-Pomeron graph exchanges)
- small number of adjustable parameters
- generalization to hA , AA - parameter free
- high energy asymptotics - governed by pQCD
- parameters can be fixed using cross sections, total and diffractive SFs, particle production

Main differences to the linear scheme:

- screening of the “soft” particle production
- in the “dense” limit (large s , small b , large A) -
re-normalization of the “soft” Pomeron intercept: $\alpha_P(0) \rightarrow \tilde{\alpha}_P(0) < 1$
 $\Rightarrow \chi_{ab}^{P_{\text{soft}}}(s, b)$ - decreasing with s - saturation at the Q_0^2 scale!
- \Rightarrow approaches “mini-jet” picture in the “dense” limit

Drawback of the scheme:

- can not treat screening & saturation effects at $Q^2 > Q_0^2$!

Air Shower Predictions

Basic EAS characteristics:

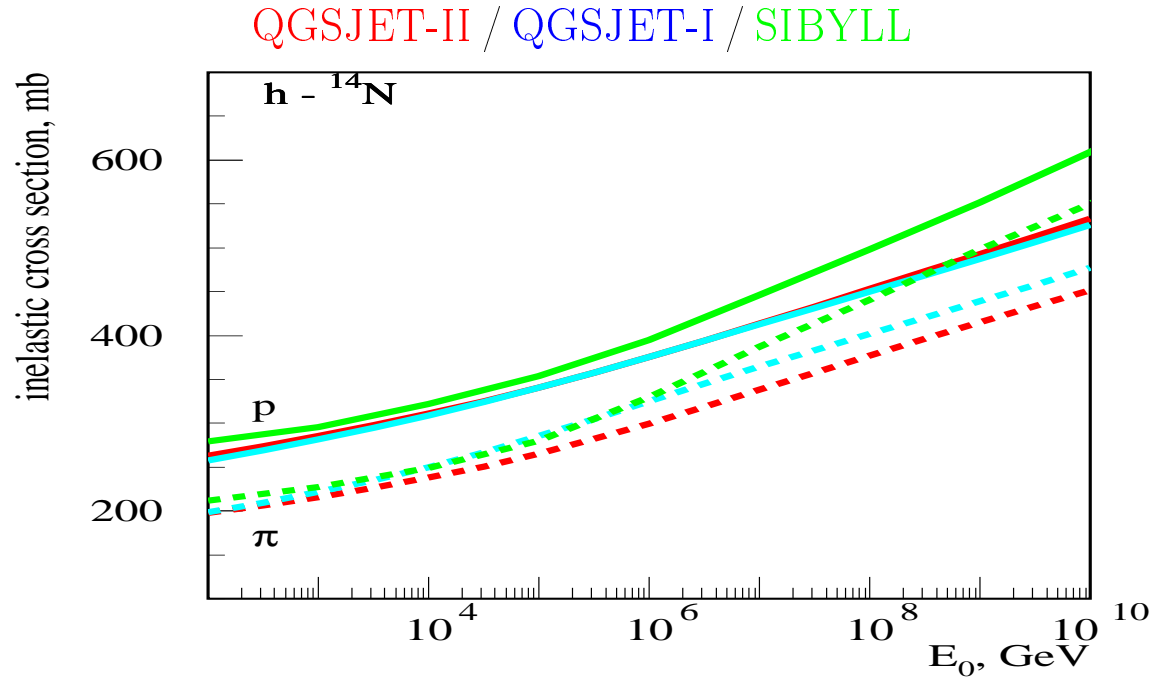
- X_{\max} - mainly defined by $\sigma_{h-\text{air}}^{\text{inel}}$, $K_{h-\text{air}}^{\text{inel}}$ (especially, $\sigma_{p-\text{air}}^{\text{inel}}$, $K_{p-\text{air}}^{\text{inel}}$)
- N_e - correlated with X_{\max}
- N_μ - strongly depends on $N_{h-\text{air}}^{\text{ch}}$ (especially, $N_{\pi-\text{air}}^{\text{ch}}$), however, also depends on the spectra

Inelastic cross sections ($\sigma_{h-\text{air}}^{\text{inel}}$) - no new parameters compared to $h - p$

- depend on Pomeron parameters / parton momentum distributions
- depend on the inelastic screening (related to the diffraction dissociation)
- depend on non-linear screening effects (stronger than for $\sigma_{h-p}^{\text{inel}}$)

$\pi - \text{air}$ / $p - \text{air}$:

- stronger inelastic screening
- stronger non-linear effects (larger relative weight, stronger parton “packing”)



$E_0 = 200 \text{ GeV}$ ($\sigma_{n-^{12}\text{C}}^{\text{inel}(\text{exp})}$ - T.Roberts et al., 1979;

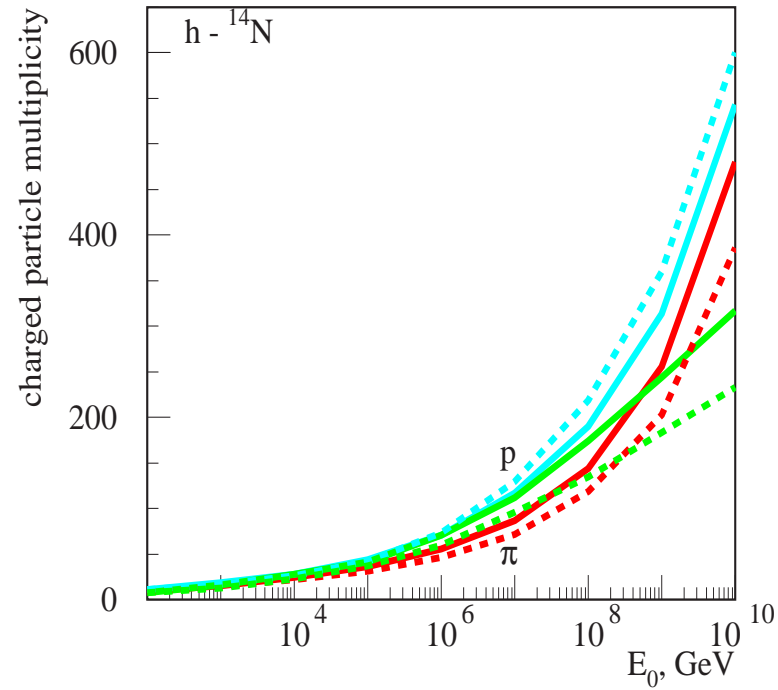
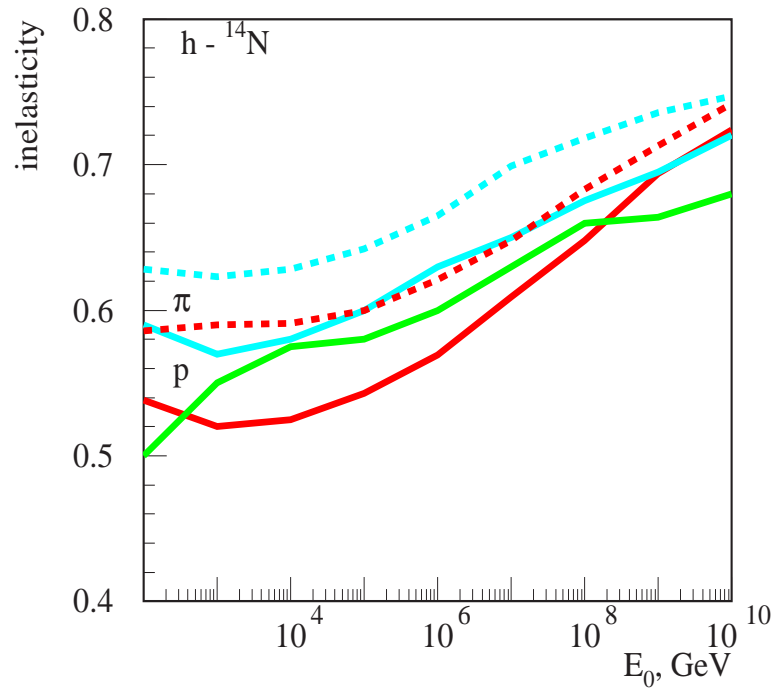
$\sigma_{h-^{12}\text{C}}^{\text{abs}(\text{exp})}$ - A.Carroll et al., 1979)

- $\sigma_{p-^{12}\text{C}}^{\text{inel}} = 241 \text{ mb}, \quad \sigma_{n-^{12}\text{C}}^{\text{inel}(\text{exp})} = 237 \pm 2 \text{ mb}$
- $\sigma_{p-^{12}\text{C}}^{\text{abs}} = 232 \text{ mb}, \quad \sigma_{p-^{12}\text{C}}^{\text{abs}(\text{exp})} = 225 \pm 7 \text{ mb}$
- $\sigma_{\pi-^{12}\text{C}}^{\text{abs}} = 170 \text{ mb}, \quad \sigma_{\pi^+-^{12}\text{C}}^{\text{abs}(\text{exp})} = 171 \pm 5 \text{ mb}$

Inelasticity ($K_{h\text{-air}}^{\text{inel}}$) and multiplicity ($N_{h\text{-air}}^{\text{ch}}$):

strongly reduced by screening effects in the “soft” component
(especially for $\pi - \text{air}$)

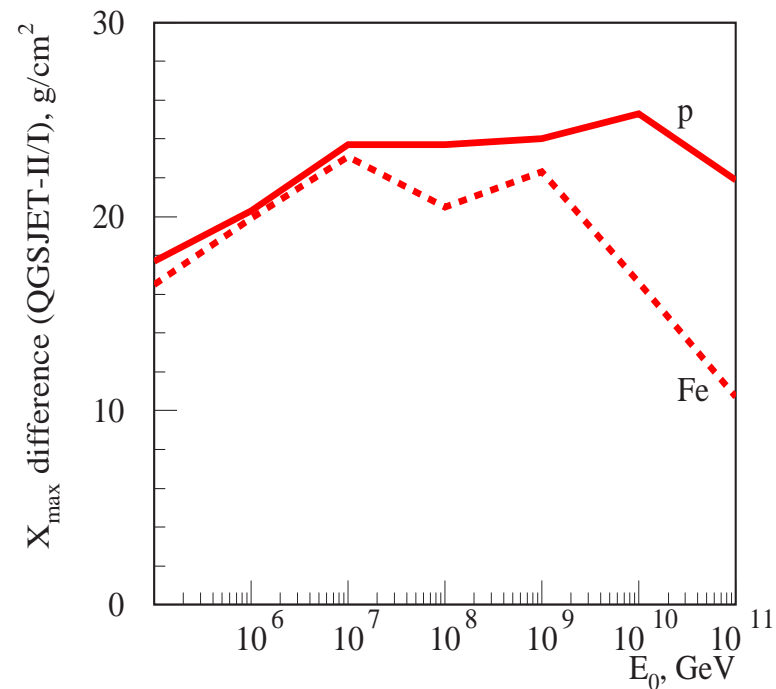
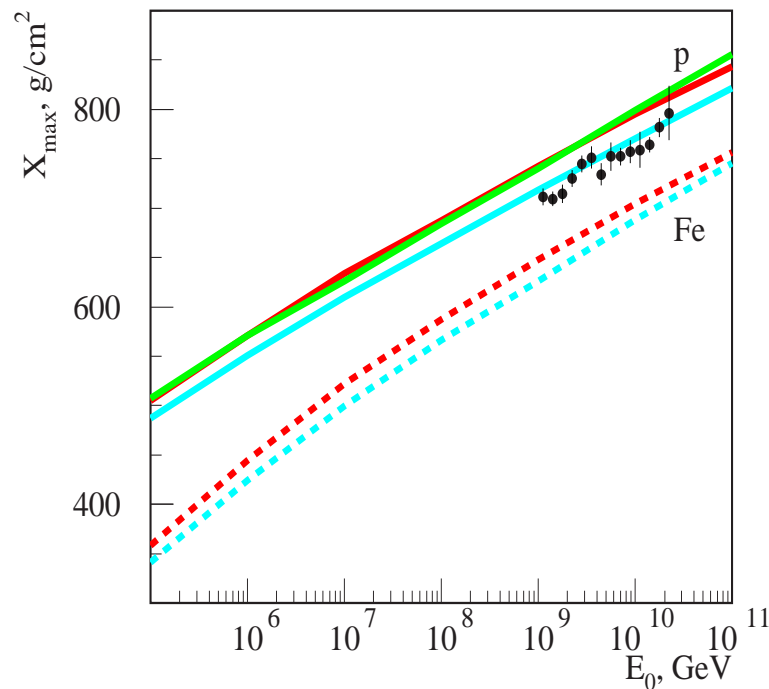
QGSJET-II / QGSJET-I / SIBYLL



Shower maximum position (X_{\max}):

- systematically shifted deeper by screening effects (by $\sim 20 \text{ g/cm}^2$)
- \Rightarrow will lead to a heavier composition
- agreement between QGSJET-II and SIBYLL

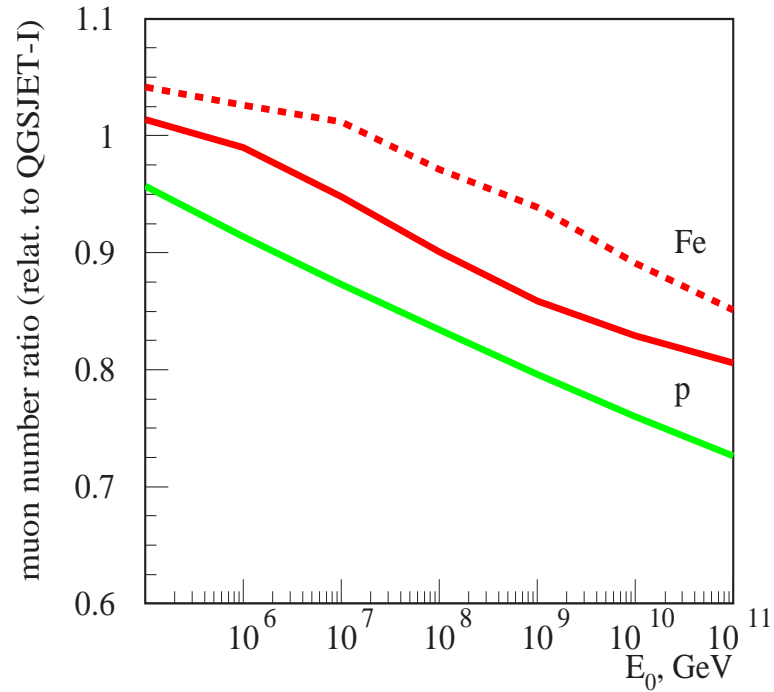
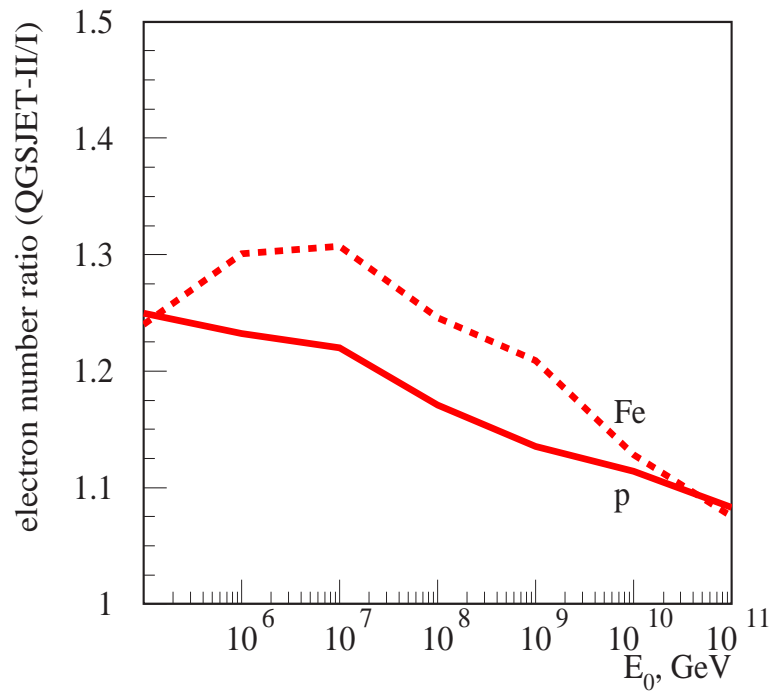
QGSJET-II / QGSJET-I / SIBYLL vrs. HIRES data (R. Abbasi et al., 2004)



Electron (N_e) and muon (N_μ , $E_\mu > 1$ GeV) number ratios:

- $N_\mu(N_e)$ - significantly increasing
- \Rightarrow again leads to a heavier composition

QGSJET-II (SIBYLL) relative to QGSJET-I

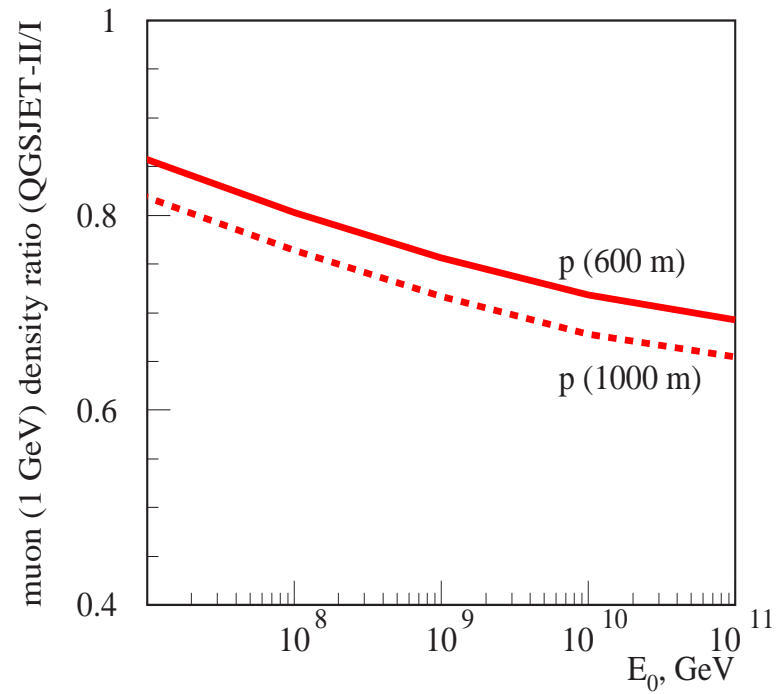
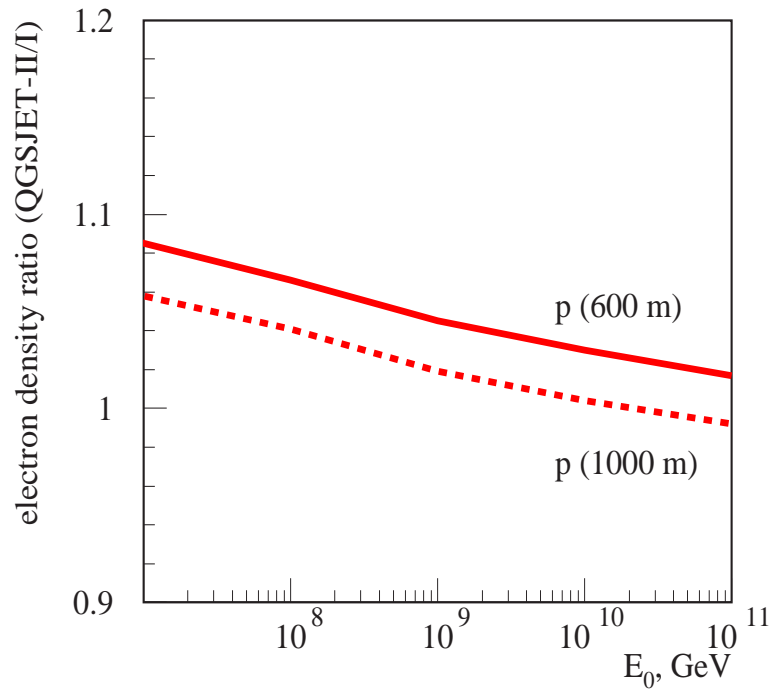


Energy determination

EAS profile: N_e^{\max} ratio - within 3%

Electron (ρ_e) and muon (ρ_μ , $E_\mu > 1$ GeV) densities at 600 and 1000 m:

- energy estimations depend on the **muon contribution** (water tanks?)



Model Uncertainties

From experimental data:

- calibration - mainly on hp , mainly at fixed target energies
- hA , AA data scarce
- central heavy ion collisions - provide only general guidance (may include new physics, e.g., QGP effects)
- $\sigma_{pp}^{\text{tot}}(\sqrt{s} = 1800 \text{ GeV})$ - two contradictory measurements
- $\sigma_{hA}^{\text{inel}}$ - only at fixed target energies

Model parameter & algorithm freedom:

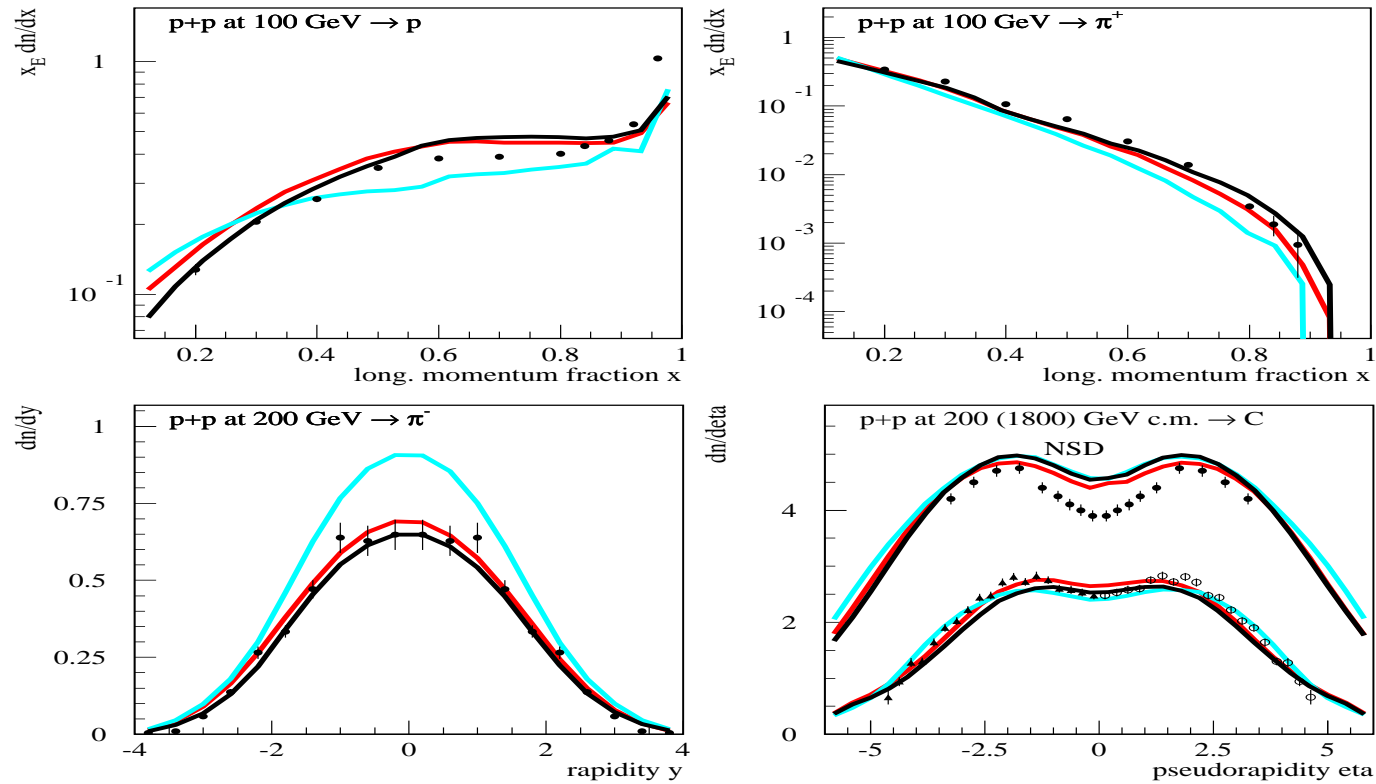
- treatment of leading states & string ends (“valence-like” or “sea-like”)
- treatment of low mass diffraction \Rightarrow inelastic screening
- choice of Q_0^2 (saturation effects)
- choice of parton SFs at the Q_0^2 scale (factorization scale, “resolved photon” contribution, etc.)
- possible loss of coherence of leading states (H.Drescher et al., 2004)

Before - tests of parameter freedom with QGSJET-I

(M.Zha, J.Knapp, SO, 2002)

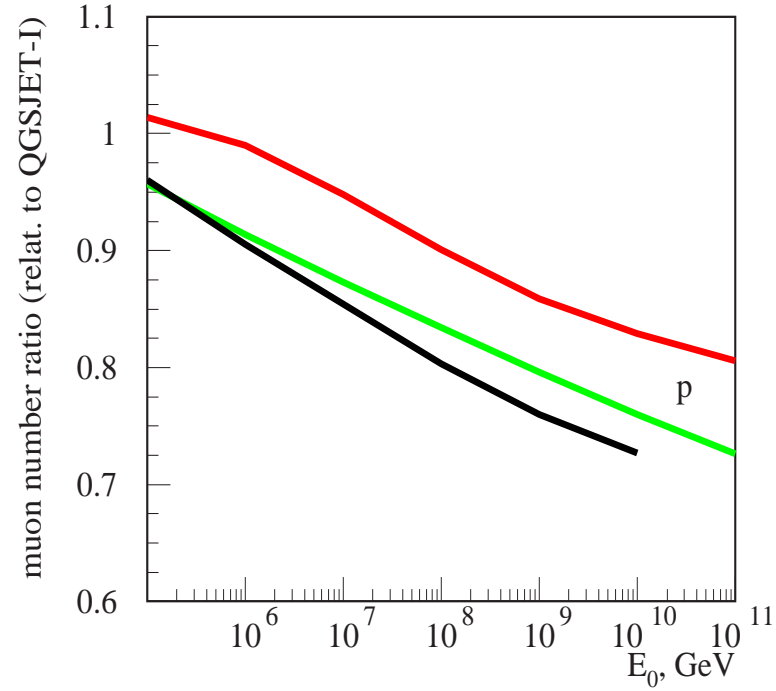
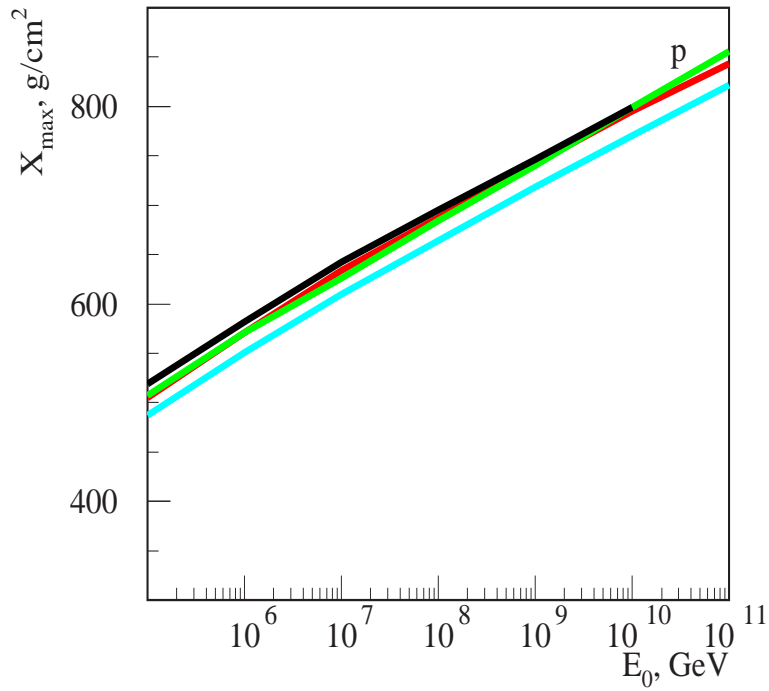
Here - only 1st option (string end momentum distribution)

QGSJET-II / QGSJET-I / QGSJET-IIm



More “sea-like” string ends \Rightarrow smaller multiplicity & inelasticity in hA (AA)
 \Rightarrow deeper X_{\max} , smaller N_{μ}

QGSJET-II / QGSJET-I / SIBYLL / QGSJET-IIIm



Future Prospects

Largely expected from QCD:

- progress with BFKL Pomeron & NLL QCD
- perturbative treatment of screening & saturation effects

⇒ reliable calculations of high energy asymptotics of scattering amplitudes

What QCD can not do (always non-perturbative):

- hadronization
- treatment of leading states

But a lot can be fixed with new experimental data:

- on particle production in hh -, hA -, AA -collisions (RHIC, LHC)
- on total and diffraction cross sections (Tevatron?, LHC)

And [present models are not bad](#):

- describe cross sections & particle production for different hadronic processes
- describe perturbative observables:
high p_t particle spectra, SFs, hard diffraction
- converge in EAS predictions

And can be further improved [on the phenomenological level](#)

Considerable progress can be made and [is made with CR data](#):

- studies of correlations of EAS observables discriminate between models
(example, KASCADE)
- measurements of $\sigma_{h-\text{air}}^{\text{inel}}$ (example, HIRES)
- penetrating hadrons - sensitive to $\sigma_{h-\text{air}}^{\text{inel}}$, $\sigma_{h-\text{air}}^{\text{difr}}$
- muon spectra & bundles - sensitive to $N_{h-\text{air}}^{\text{ch}}$, hadron spectra

[We made a big progress!](#)
[Things will further improve in future!](#)